

ATMOSPHERIC DRAG MODELLING AND FORECASTING

MISSION ANALYSIS & OPERATIONS

2022/05/13

COMET Space Weather & Orbitography

Nicolas TCHINTCHARADZÉ / Hugo VEUILLEZ

CNES



Introduction

2016-2020 summary

Go forward ?

Partial derivatives

A large, light blue, stylized number '1' is positioned on the left side of the slide, partially cut off by the edge.

INTRODUCTION

Introduction

What matters?

- ❖ **On LEO orbits**, the orbit prediction is imperfect :
 - all contributions to drag are with **uncertainties**

$$\gamma = C_f \frac{1}{2} \rho \frac{SC_x}{m} V^2$$

2015 expectations

- ❖ **Characterize** the orbital prediction errors in **MISSION ANALYSIS**
 - Short term : a few hours
 - Mid-term : a few days
 - Long term : N x 10 days
- ❖ **Improve** the **operational** process of orbital prediction
 - Choice of atmospheric model
 - Choice of solar activity inputs
 - Choice of S.Cx model
 - Choice of strategy : how to apply the observed error?

Goal of this presentation?
Share our state of art

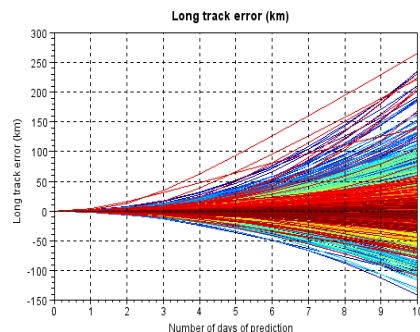
2016 state of art – by examples

Additive method for mission analysis

- ❖ Operational period expected? choice of a past observed period
predicted and real flux/indices usage
- ❖ Mean surface and constant Cx

Cf. CCT ORB 10/04/2014

Using a dedicated tool, we obtain the error due to solar activity uncertainty



Error >0 ? drag *under* estimated

Reality minus Prediction

↔ *density model with observed actsol*
minus density model with predicted actsol

Error <0 ? drag *over* estimated

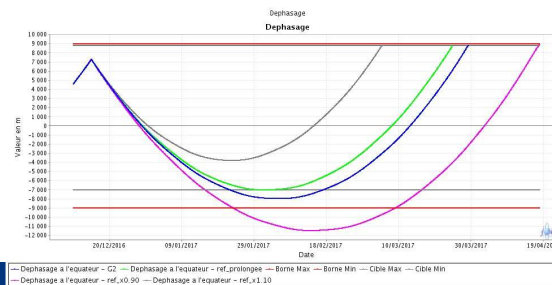
Fine .. But ...

- Only absolute error analysis
- Density model not challenged
- Too oversizing ?
- Some observed cases of undersizing ..

- ❖ And we added a 20% margin due to observation on Cf

Operational problematic

- ❖ *Unknown future*



Empiric approach

- Limited trust in operational tools ...



2016-2020 SUMMARY

MODELS, TOOLS

2016 workplan

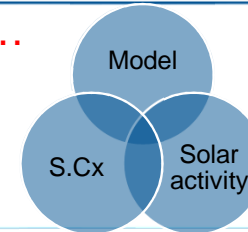
Observe

- Use of available operational data ELISA, SPOT, HELIOS, JASON, GRACE, ...
 - Error between prediction and observation, with different models, different data : F10.7 and F30 (R&T)
 - Operational Cf restituted
 - Relative errors instead of absolute errors
 - S.Cx model tests

Analyze

- Working group

Some lack of time / wrong directions ...
4 correlated thematics !



Improve

- Empiric method for mission analysis? The initial driver
- Preconisations for operations? The final purpose
- Monitoring in operations ? The dream
- Go forward? The 2022 status-quo

Thematic#1 : atmospheric model



Several comparisons to « identify » which model « fits the reality »

1. On 3 extrapolation duration (1d, 10d, 30d)
2. With observed solar activities

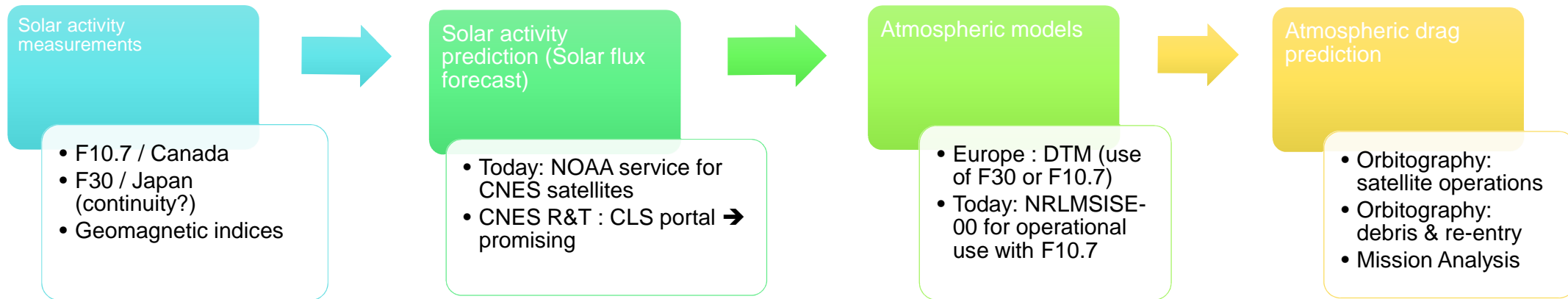
For : DM2013 F10.7 / DM2013 F30 / MSIS2000 / Jacchia 2008

Bias identified : GS recommandation to check the « meta model » of S.Cx

MSIS2000 allways used in operational tools

But DTM 2013 F30 to target (and SWAMI at mid term)

Thematic#2 : Solar activity forecast



- ❖ analyses show better drag predictions with F30 than with F10.7 (through the DTM model)
- ❖ CNES development and investment (R&T) in F30 and F10.7 predictions together with CLS and LPC2E support (CLS: <https://spaceweather.cls.fr/services/radioflux/>)

Thematic#3 : S.Cx



Elaboration with Cook formula of « 3D » S.Cx model for a spacecraft (CERES)

- ❖ **PERT tool : S.Cx table as a function of altitude and solar activity**
- ❖ **PATRIUS tool : Global Drag force interpolating this table at each instant**

Operational use since March 2022

Thematic#4 : mission analysis method (1/2)

Empiric intermediate method

$$\gamma = \frac{1}{2} \rho \frac{S C_x}{m} V^2$$

$XX_P = \text{Predicted}$

$XX_R = \text{Real and Unknown}$

$XX_{mod} = \text{Computed with observed flux}$

$$\text{Drag error : } \varepsilon_F = \gamma_P - \gamma_R = \left[1 - \frac{(SC_x)_R}{(SC_x)_P} \frac{\rho_R}{\rho_{mod}} \frac{\rho_{mod}}{\rho_P} \right] \gamma_P$$

$$\text{Along track error : } \varepsilon_{LT} \cong \iint_t 3\varepsilon_F dt^2 = \left[1 - \tau_{SC_x} \tau_{mod} \tau_{actsol} \right] \iint_t 3\gamma_P dt^2$$

LT_P : along track predicted distance

$$\text{Relative error : } \tau_{total} = \tau_{SC_x} \tau_{mod} \tau_{actsol} = 1 - \frac{\varepsilon_{LT}}{LT_P}$$

Model error



was managed

via the previous 20% margin due to observation on Cf

Solar activity error

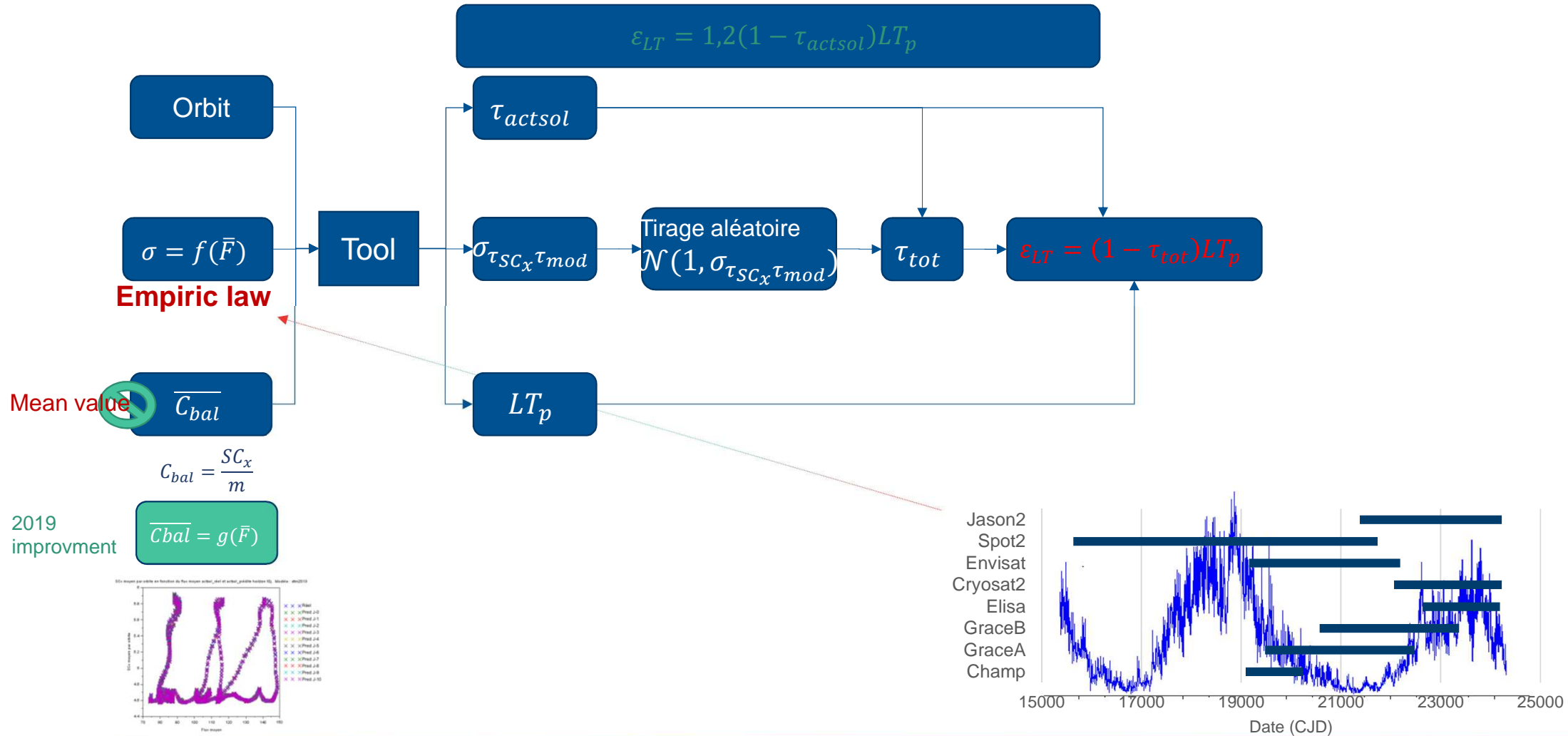


managed

by the 2016 existing tool

Thematic#4 : mission analysis method (2/2)

Previous method was without $\tau_{SCx} \tau_{mod} \dots$



GO FORWARD

Need of a new start

Mission analysis point of view

- ❖ The actual empiric method is not perfect, but is **our actual reference**

Operational point of view

- ❖ Cf prediction usage workplan?
- ❖ Partial derivatives approach (see last section)
- ❖ DTM2013 or SWAMI deployment to anticipate

Other Ideas

- ❖ With LPC2E/CLS help : F30 forecast consolidation (Europe independance, ...)
- ❖ Prediction of geomagnetic indices
- ❖ Long term dream? develop « SpaceWeather laboratory »
 - use the daily CNES data,
 - continue to explore,
 - Machine learning ...
- ❖ SWAMI usage

4

CURRENT AND FUTUR WORKING AXES

The drag acceleration and the sources of uncertainty

We can represent the **drag acceleration** with the following equation:

$$\gamma = C_D * \frac{1}{2} \rho \frac{SC_x}{m} V^2$$

- **Low contribution to uncertainty:**

- **Velocity:** a few meters per second (spacecraft + atmosphere velocity) of uncertainty relative to several kilometers per second
- **Mass:** a few grams compared to hundreds of kilograms.

- **High contribution to uncertainty:**

- SC_x : the reference area and the drag coefficient imply complex physics. They mainly depend on the *attitude* of the spacecraft. We can introduce the following notations:

$$SC_x(Attitude) = \widehat{SC}_x(Attitude) + \varepsilon_{SC_x}(Attitude)$$

Notations

X : Real value
 \hat{X} : Model value
 ε : Error

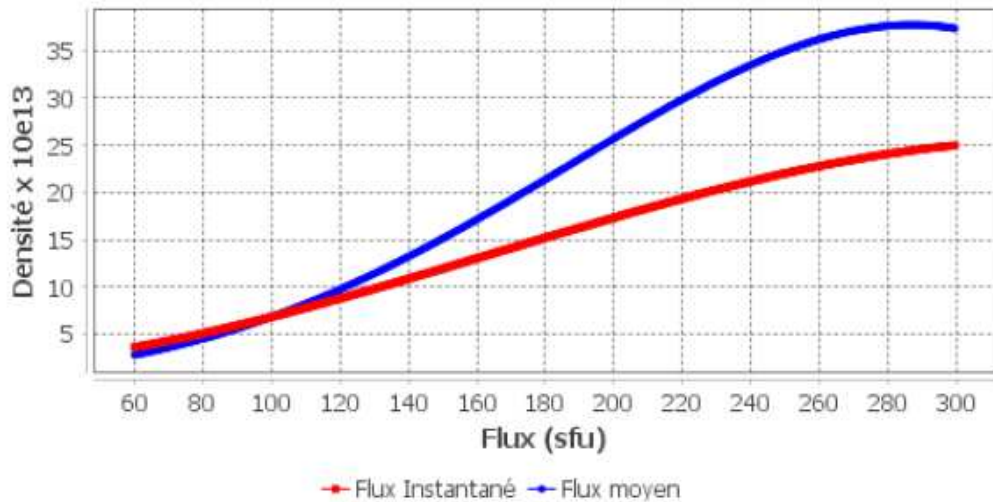
- **Atmospheric density:** implies complex physics of the whole earth system. It has an important dependence to the *solar activity*. We can introduce the following notations:

$$\rho = \hat{\rho}(SolAct) + \epsilon_\rho$$

- C_D : adjustment coefficient to fit observations. It *absorbs* all the *lack of modelling* of the previous bullets.

Atmospheric model partial derivatives

The errors of the atmosphere density are the addition of:

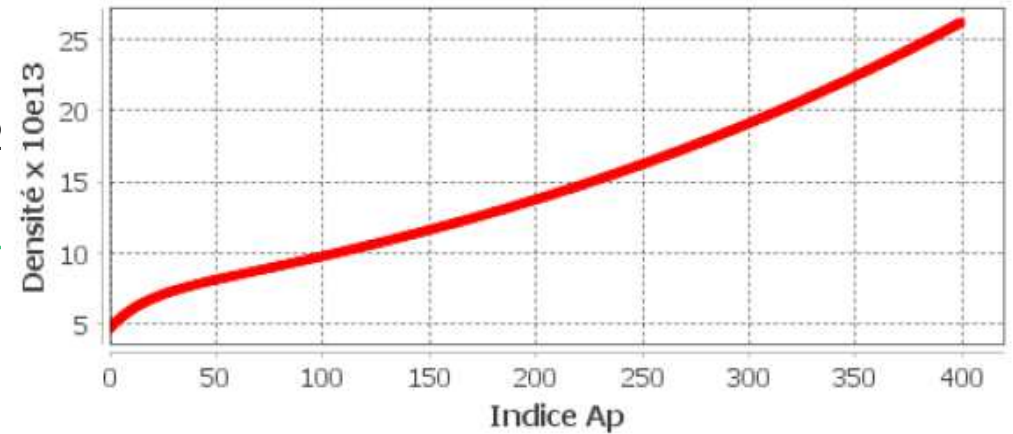


Notations

X : Real value

$\hat{X}(t + \Delta t)$: Model value at date $t + \Delta t$

$\varepsilon(t, \Delta t)$: Model error at date $t + \Delta t$. depends on

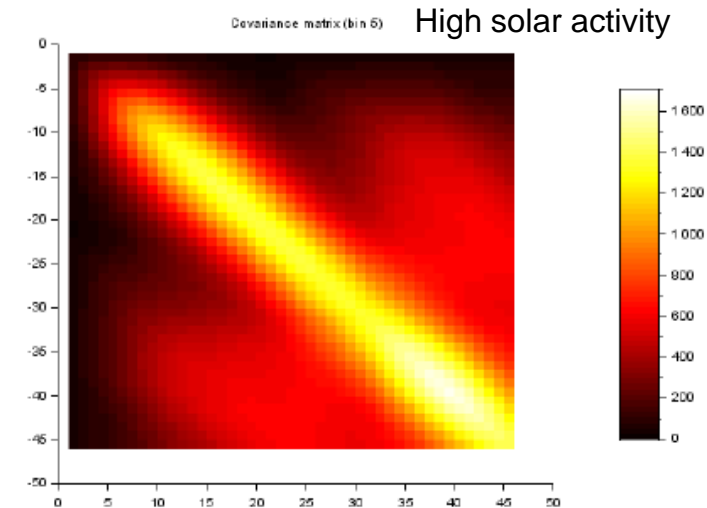
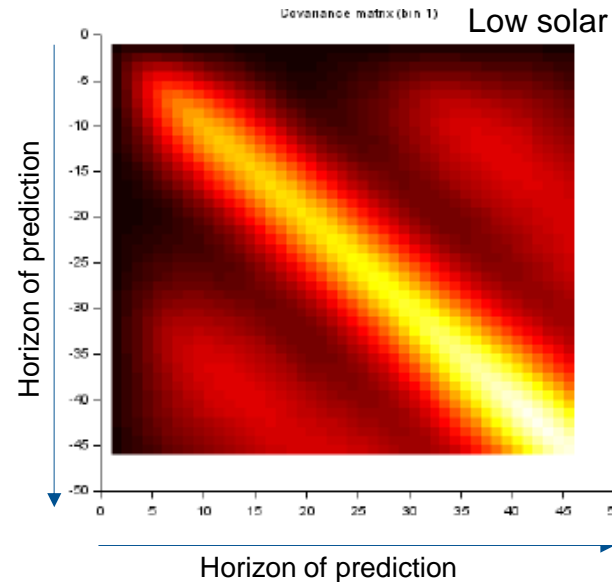
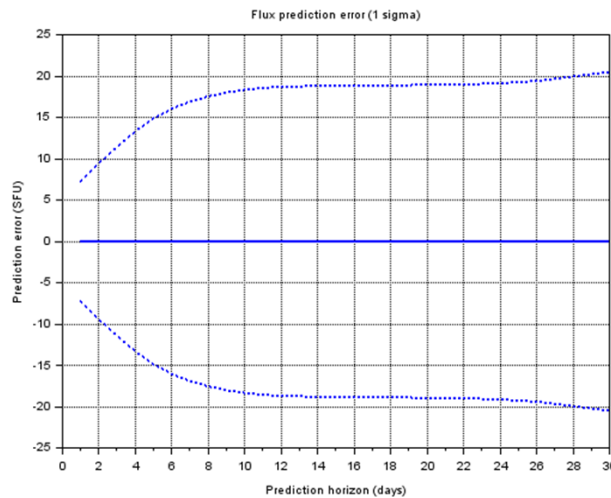


- These partial derivatives of the atmospheric model allow to compute trajectory partial derivatives (through numerical integration):

$$dX \approx \frac{\partial X}{\partial Fl} dFl + \frac{\partial X}{\partial Ap} dAp$$

Atmospheric model partial derivatives

- Good knowledge of the solar activity error statistics thanks to a long historic of data
 - Warning: the **temporal correlation** between predictions must be taken into account



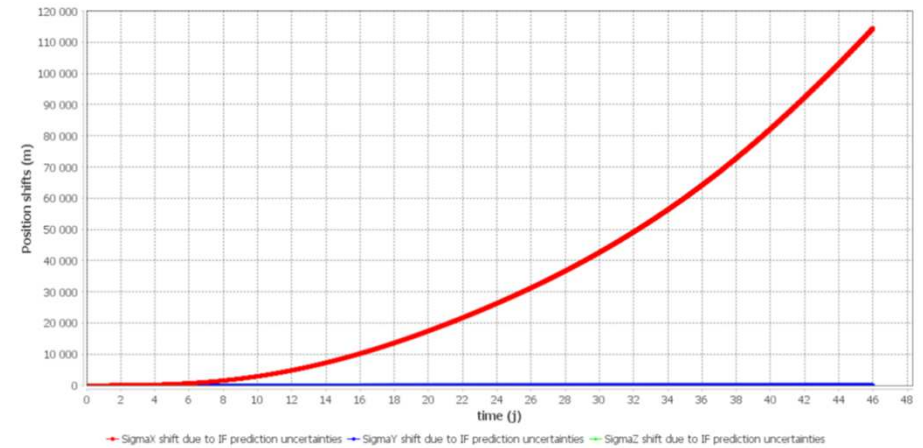
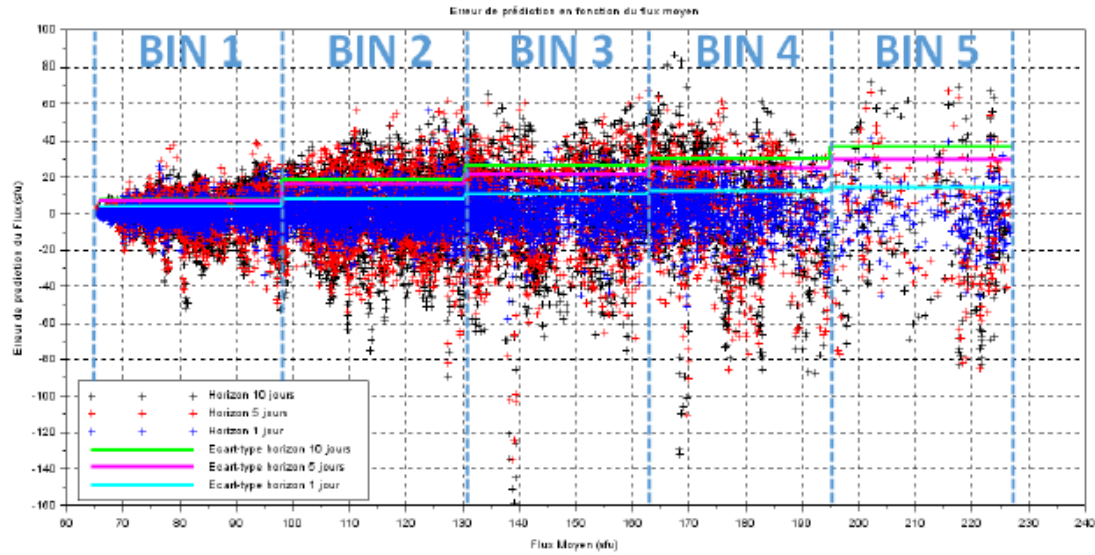
- Thus, for a given orbit, we can have a good knowledge of the orbital prediction error statistics due to the solar activity prediction errors
- How to use this info?
 - To better understand the contribution of each post of error
 - For mission analysis: Take into account a better coupling between orbit determination errors and orbit prediction errors
 - For covariance propagation
 - **Difficulty:** How to integrate the other drag acceleration post of errors to have a full error budget ?

Futur work



- Include the draft of the partial derivatives computation in our softwares
- Define a protocol to use it in operations for covariance prediction (collision risk management)
- Work on the best way to use the Cd estimation for the orbit predictions

Annex: orbit error from solar activity prediction errors



Bin représenté	1	2	3	4	5
Ecart-type de la dérive (en km)	1,6	9,5	37	90	114

After 45 days of prediction